

The complex variability of blazars: Time-scales and periodicity analysis in S4 0954+65

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Abstract

Among active galactic nuclei, blazars show extreme variability properties. We here investigate the case of the BL Lac object S4 0954+65 with data acquired in 2019-2020 by the Transiting Exoplanet Survey Satellite (TESS) and by the Whole Earth Blazar Telescope (WEBT) Collaboration. The 2-min cadence optical light curves provided by TESS during three observing sectors of nearly 1 month each allow us to study the fast variability in great detail. We identify several characteristic short-term time-scales, ranging from a few hours to a few days. However, these are not persistent, as they differ in the various TESS sectors. The long-term photometric and polarimetric optical and radio monitoring undertaken by the WEBT brings significant additional information, revealing that (i) in the optical, long-term flux changes are almost achromatic, while the short-term ones are strongly chromatic; (ii) the radio flux variations at 37 GHz follow those in the optical with a delay of about 3 weeks; (iii) the range of variation of the polarization degree and angle is much larger in the optical than in the radio band, but the mean polarization angles are similar; (iv) the optical long-term variability is characterized by a quasi-periodicity of about 1 month. We explain the source behaviour in terms of a rotating inhomogeneous helical jet, whose pitch angle can change in time.

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Keywords

BL Lacertae objects: general, BL Lacertae objects: individual: S4 0954+65, Galaxies: active, Galaxies: jets

References

- [1] Ackermann M. et al., 2015, *ApJ*, 813, L41
- [2] Agarwal A. et al., 2015, *MNRAS*, 451, 3882
- [3] Agudo I., Thum C., Wiesemeyer H., Krichbaum T. P., 2010, *ApJS*, 189, 1
- [4] Agudo I., Thum C., Gómez J. L., Wiesemeyer H., 2014, *A&A*, 566, A59
- [5] Agudo I., Thum C., Ramakrishnan V., Molina S. N., Casadio C., Gómez J. L., 2018a, *MNRAS*, 473, 1850
- [6] Agudo I. et al., 2018b, *MNRAS*, 474, 1427
- [7] Bach U. et al., 2006, *A&A*, 456, 105
- [8] Bachev R., 2015, *MNRAS*, 451, L21
- [9] Becerra González J., Acosta-Pulido J. A., Boschin W., Clavero R., Otero-Santos J., Carballo-Bello J. A., Domínguez-Palmero L., 2021, *MNRAS*, preprint (arXiv:2010.14532)
- [10] Bessell M. S., Castelli F., Plez B., 1998, *A&A*, 333, 231
- [11] Blinov D. et al., 2015, *MNRAS*, 453, 1669
- [12] Blinov D. et al., 2016, *MNRAS*, 462, 1775
- [13] Blinov D. et al., 2018, *MNRAS*, 474, 1296
- [14] Britzen S., Witzel A., Krichbaum T. P., Beckert T., Campbell R. M., Schalinski C., Campbell J., 2005, *MNRAS*, 362, 966
- [15] Britzen S. et al., 2017, *A&A*, 602, A29
- [16] Carnerero M. I. et al., 2015, *MNRAS*, 450, 2677
- [17] Carnerero M. I. et al., 2017, *MNRAS*, 472, 3789
- [18] Casadio C. et al., 2015, *ApJ*, 813, 51
- [19] Cheng X.-L., Zhang Y.-H., Xu L., 2013, *MNRAS*, 429, 2773
- [20] Conway J. E., Wrobel J. M., 1995, *ApJ*, 439, 98
- [21] Deeming T. J., 1975, *Ap&SS*, 36, 137
- [22] Edelson R. A., Krolik J. H., 1988, *ApJ*, 333, 646
- [23] Fromm C. M., Ros E., Perucho M., Savolainen T., Mimica P., Kadler M., Lobanov A. P., Zensus J. A., 2013, *A&A*, 557, A105
- [24] Giroletti M., Righini S., 2020, *MNRAS*, 492, 2807
- [25] Gupta A. C. et al., 2017, *MNRAS*, 472, 788
- [26] Hagen-Thorn V. A. et al., 2015, *Astron. Rep.*, 59, 551
- [27] Heidt J., Wagner S. J., 1996, *A&A*, 305, 42
- [28] Horne J. H., Baliunas S. L., 1986, *ApJ*, 302, 757
- [29] Hufnagel B. R., Bregman J. N., 1992, *ApJ*, 386, 473
- [30] Jorstad S. G. et al., 2017, *ApJ*, 846, 98
- [31] Lainela M. et al., 1999, *ApJ*, 521, 561
- [32] Landoni M., Falomo R., Treves A., Scarpa R., Reverte Payá D., 2015, *AJ*, 150, 181
- [33] Larionov V. M., Villata M., Raiteri C. M., 2010, *A&A*, 510, A93
- [34] Larionov V. M. et al., 2013, *ApJ*, 768, 40
- [35] Larionov V. M. et al., 2016, *MNRAS*, 461, 3047
- [36] Larionov V. M. et al., 2020, *MNRAS*, 492, 3829
- [37] Lawrence C. R., Pearson T. J., Readhead A. C. S., Unwin S. C., 1986, *AJ*, 91, 494
- [38] Liska M., Hesp C., Tchekhovskoy A., Ingram A., van der Klis M., Markoff S., 2018, *MNRAS*, 474, L81
- [39] Lyutikov M., Kravchenko E. V., 2017, *MNRAS*, 467, 3876
- [40] MAGIC Collaboration, 2018, *A&A*, 617, A30
- [41] Marchili N., Krichbaum T. P., Liu X., Song H. G., Gabányi K. É., Fuhrmann L., Witzel A., Zensus J. A., 2012, *A&A*, 542, A121
- [42] Marscher A. P., 2014, *ApJ*, 780, 87
- [43] Marscher A. P. et al., 2008, *Nature*, 452, 966
- [44] McHardy I.M., Marscher A. P., GearW. K., MuxlowT., Lehto H. J., Abraham R. G., 1990, *MNRAS*, 246, 305
- [45] Mignone A., Rossi P., Bodo G., Ferrari A., Massaglia S., 2010, *MNRAS*, 402, 7
- [46] Moll R., Spruit H. C., Obergaulinger M., 2008, *A&A*, 492, 621
- [47] Morozova D. A. et al., 2014, *AJ*, 148, 42
- [48] Nakamura M., Uchida Y., Hirose S., 2001, *New A*, 6, 61

- [49] Ostorero L., Villata M., Raiteri C. M., 2004, A&A, 419, 913
- [50] Otero-Santos J. et al., 2020, MNRAS, 492, 5524
- [51] PeruchoM., Kovalev Y. Y., Lobanov A. P., Hardee P. E., Agudo I., 2012, ApJ, 749, 55
- [52] Peterson B. M., Wanders I., Horne K., Collier S., Alexander T., Kaspi S., Maoz D., 1998, PASP, 110, 660
- [53] Raiteri C. M. et al., 1999, A&A, 352, 19
- [54] Raiteri C. M. et al., 2001, A&A, 377, 396
- [55] Raiteri C. M. et al., 2003, A&A, 402, 151
- [56] Raiteri C. M. et al., 2009, A&A, 507, 769
- [57] Raiteri C. M. et al., 2013, MNRAS, 436, 1530
- [58] Raiteri C. M. et al., 2015, MNRAS, 454, 353
- [59] Raiteri C. M. et al., 2017a, MNRAS, 466, 3762
- [60] Raiteri C. M. et al., 2017b, Nature, 552, 374
- [61] Raiteri C. M. et al., 2021, MNRAS, 501, 1100
- [62] Ricker G. R. et al., 2015, J. Astron. Telesc. Instrum. Syst., 1, 014003
- [63] Rieger F. M., 2004, ApJ, 615, L5
- [64] Sandrinelli A. et al., 2017, A&A, 600, A132
- [65] Scargle J. D., 1982, ApJ, 263, 835
- [66] Sillanpaa A., Haarala S., Valtonen M. J., Sundelius B., Byrd G. G., 1988, ApJ, 325, 628
- [67] Sillanpaa A. et al., 1996, A&A, 315, L13
- [68] Simonetti J. H., Cordes J. M., Heeschen D. S., 1985, ApJ, 296, 46
- [69] Smith P. S., 1996, in Miller H. R., Webb J. R., Noble J. C., eds, ASP Conf. Ser. Vol. 110, Blazar ContinuumVariability.Astron. Soc. Pac., San Francisco, p. 135
- [70] Sobacchi E., Sormani M. C., Stamerra A., 2017, MNRAS, 465, 161
- [71] Stickel M., Fried J. W., Kuehr H., 1993, A&AS, 98, 393
- [72] Teräsranta H. et al., 1998, A&AS, 132, 305
- [73] Thum C., Agudo I., Molina S. N., Casadio C., Gómez J. L., Morris D., Ramakrishnan V., Sievers A., 2018, MNRAS, 473, 2506
- [74] Urry C. M., Padovani P., 1995, PASP, 107, 803
- [75] Valtonen M., Pihajoki P., 2013, A&A, 557, A28
- [76] Vaughan S., 2005, A&A, 431, 391
- [77] Vaughan S., Edelson R., Warwick R. S., Uttley P., 2003, MNRAS, 345, 1271
- [78] Villata M., Ferrari A., 1994a, Phys. Plasmas, 1, 2200
- [79] Villata M., Ferrari A., 1994b, A&A, 284, 663
- [80] Villata M., Ferrari A., 1995, A&A, 293, 626
- [81] Villata M., Raiteri C. M., 1999, A&A, 347, 30
- [82] Villata M., Tsinganos K., 1993, Phys. Fluids B, 5, 2153
- [83] Villata M., Tsinganos K., 1994, Phys. Plasmas, 1, 216
- [84] Villata M., Raiteri C. M., Sillanpaa A., Takalo L. O., 1998, MNRAS, 293, L13
- [85] Villata M. et al., 2002, A&A, 390, 407
- [86] Villata M. et al., 2004, A&A, 421, 103
- [87] Villata M. et al., 2006, A&A, 453, 817
- [88] Villata M. et al., 2008, A&A, 481, L79
- [89] Wagner S., Sanchez-Pons F., Quirrenbach A., Witzel A., 1990, A&A, 235, L1
- [90] Wagner S. J. et al., 1993, A&A, 271, 344
- [91] Weaver Z. R. et al., 2020, ApJ, 900, 137
- [92] Wu J., Peng B., Zhou X., Ma J., Jiang Z., Chen J., 2005, AJ, 129, 1818
- [93] Zhang H., Li X., Giannios D., Guo F., Liu Y.-H., Dong L., 2020, ApJ, 901, 149